

Physics 114 Statistical Mechanics Spring 2021  
Week 6 Conceptual Overview

Concept checklist from Readings:

- *About entropy of mixing:* G&T section 6.1 show us by explicitly using  $S(T, V, N)$ , that mixing identical gas particles results in  $\Delta S = 0$ , as it should be :-)
- *Applications of the Boltzmann distribution:* It can be applied, as in G&T section 6.2.2 and 6.2.3 to gas velocities. In that case, it is often known as the *Maxwell-Boltzmann (MB) distribution* and the speed distribution  $p \propto v^2 e^{-mv^2/2kT}$  is called a *Maxwellian*. Ch. 5 of B&B is a short, excellent reference for the MB distribution of velocities and speeds.
- A second application of the Boltzmann distribution is the beautiful *equipartition theorem*. This is done in G&T section 6.2.1 and B&B Ch. 19. The theorem can be summarized by saying that if a particle has  $f$  independent, quadratic terms in its Hamiltonian, then its expected equilibrium energy is  $\langle E \rangle = \frac{f}{2} kT$ .
- A natural consequence for gasses is that  $C_v = \frac{f}{2} R$  per mole. Similarly,  $C_p = (\frac{f}{2} + 1) R$  and the adiabatic constant is  $\gamma = 1 + \frac{2}{f}$ .
- Equipartition is derived assuming a continuous distribution of energies. Thus, if  $kT \gg \Delta E_q$ , where  $\Delta E_q$  is a typical separation between quantum levels, equipartition applies. On the other hand, if  $kT \ll E_1 - E_0$ , that mode will be “stuck” in its ground state whose energy is  $E_0$ .
- B&B have several examples of equipartition: a harmonic oscillator, a solid (modeled as a set of  $N$  harmonic oscillators in  $3d$ , monatomic and diatomic gasses). Be sure that you can use equipartition to find their average energies and their specific heats.
- $Z(T, V, N)$  is the partition function for  $N$  particles, but we read a lot about the *single particle partition function*. Why care about one particle? Because  $Z(T, V, 1)$  is often much easier to calculate, and if we have  $N$  distinguishable particles, it is the case that

$$Z(T, V, N) = Z(T, V, 1)^N \text{ distinguishable particles}$$

- In a related way, if we have several completely different degrees of freedom (like A = rotations, B = translations, and C = vibrations) then partition functions just multiply

$$Z = Z_A Z_B Z_C$$

- What about  $N$  indistinguishable particles?
  - The *right way*: Find out if they are bosons and fermions and do the correct counting arguments
  - The *not-totally-right way*: This works for classical ideal gasses, and is essential to get the entropy of mixing correct:

$$Z(T, V, N) = \frac{1}{N!} Z(T, V, 1)^N \quad \text{dilute, semiclassical particles}$$

This approach is predicated on the idea that there are so many states, and so few particles ( $N$  is few? Yes,  $N$  is few compared to the number of available quantum states.) that it would be extremely unlikely to find any two particles occupying the same quantum state.

- For semiclassical particles (now we are being more general than just talking about ideal gasses) G&T Section 6.2 tells us that we can cut through all the “particle in a box” counting. There is one quantum state per “box” of size  $dx dp/h$ . To find  $Z$  amounts to taking an integral over  $x_1, y_1, z_1, x_2, \dots, p_{x,1}, p_{y,1}, \dots$ . We integrate over the  $3N$  coordinates of  $N$  particles in  $3d$ . Thus

$$Z_{N, \text{classical}} = \frac{1}{h^{3N} N!} \int e^{-\beta E(x_1, y_1, \dots, p_{y,N}, p_{z,N})} dx_1, \dots, p_{z,N}$$

- How do we know a gas can be treated semi classically? For this, we calculate the *thermal deBroglie wavelength*,  $\bar{\lambda}$ . If this is much smaller than the mean distance between particles, the semiclassical limit is good.
- **From  $Z$  comes the free energy, and more:** Last week in B&B problem 14.8 we saw that

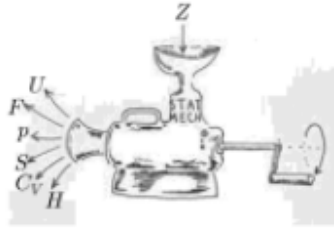
$$S/k = \beta \bar{E} + \ln Z$$

(except that B&B used  $U$ , not  $\bar{E}$  for energy.) This week, this equation reappears; moreover we learn that Helmholtz free energy is the “natural” free energy for the Canonical ensemble:

$$F(T, V, N) = -kT \ln Z(T, V, N) \quad \text{or equivalently} \quad Z = e^{-\beta F}$$

As the logarithm of a “master function”, free energy is also a master function. Either produces all thermodynamic information.

- The partition function as a sausage-making machine:



produces, via explicit formulae we already know, all useful functions of state.

Function of state		Statistical mechanical expression
$U$		$-\frac{d \ln Z}{d\beta}$
$F$		$-k_B T \ln Z$
$S$	$= -\left(\frac{\partial F}{\partial T}\right)_V = \frac{U-F}{T}$	$k_B \ln Z + k_B T \left(\frac{\partial \ln Z}{\partial T}\right)_V$
$p$	$= -\left(\frac{\partial F}{\partial V}\right)_T$	$k_B T \left(\frac{\partial \ln Z}{\partial V}\right)_T$
$H$	$= U + pV$	$k_B T \left[ T \left(\frac{\partial \ln Z}{\partial T}\right)_V + V \left(\frac{\partial \ln Z}{\partial V}\right)_T \right]$
$G$	$= F + pV = H - TS$	$k_B T \left[ -\ln Z + V \left(\frac{\partial \ln Z}{\partial V}\right)_T \right]$
$C_V$	$= \left(\frac{\partial U}{\partial T}\right)_V$	$k_B T \left[ 2 \left(\frac{\partial \ln Z}{\partial T}\right)_V + T \left(\frac{\partial^2 \ln Z}{\partial T^2}\right)_V \right]$

**Table 20.1** Thermodynamic quantities derived from the partition function  $Z$ .

- Analogy:  $\Omega(E, V, N)$  and its logarithm  $S(E, V, N)$  contain all thermodynamic information in the micro canonical ensemble.
- The protocol for the Canonical ensemble is simply described in a box in Ch. 20 of B&B
  1. Write down  $Z$
  2. Go through standard procedures to obtain functions of state you want from  $Z$
- ***Applications of Canonical ensemble to single-particle systems, with easy extension to  $N$  particles:*** Often partition functions for  $N$  particles are simply derived from 1 particle ones. This is possible if the particles don't interact! This explains why many of the examples this week are pitched in terms of finding  $f = -kT \ln Z(T, V, 1)$ ,  $s = -(\partial f / \partial T)_V$ ,  $\bar{e} = -\partial \ln Z(T, V, 1) / \partial \beta$ , ... and so on. These are the free energy, entropy, and expected energy *per particle*. For  $N$  non-interacting particles, you just *multiply by  $N$* .
- Many examples are done this week, which are good “toy” models for realistic systems. Please be sure that for the models listed below, you feel comfortable finding both single-particle and  $N$ -particle quantities like  $\bar{E}$ ,  $S$ ,  $C$ , ...
  - 2-level system
  - $N$ -level system
  - 1d harmonic oscillator
  - Rotating diatomic molecule
  - Chemical reactions: dependence of products on temperature
  - The isothermal atmosphere
  - The spin 1/2 paramagnet
  - The ideal gas
- The specific heat  $C = \partial E / \partial T$  is of particular interest this week. Please be sure you can see a calculation through, beginning with  $Z$  and ending with a “response function” like specific heat.
- We have confidence that the *equipartition theorem* holds for sufficiently high  $T$ , and this gives us a guide to what  $\bar{E}$  and  $C$  should be in the high temperature limit.
- *Magnetic systems* have a couple of observables that we care about in addition to  $\bar{E}$  and  $C$ . These are the magnetization  $m = -(\partial F / \partial B)_T$  and the susceptibility  $\chi \propto \partial m / \partial B$ . (It is the case that  $\chi \propto m/B$  in the

limit that  $B \rightarrow 0$ .) We will do more with paramagnets in a later week; for now just appreciate that we can find these things in the Canonical ensemble :-)

- ***The Canonical ensemble for ideal gasses:*** Ch. 22 of B&B is all about the ideal gas; a review and extension of what we did last week. In particular it

- reviews finding the density of states  $g(k)$  and  $g(E)$  with particle-in-a-box counting arguments.
- emphasizes the importance of the thermal de Broglie wavelength, now called  $\lambda_{th}$ . The *quantum concentration* is  $n_Q = 1/\lambda_{th}^3$ . When particles are much less concentrated than  $n_Q$ , they can be treated classically. An elegant way to write the single particle partition function for the ideal gas is:

$$Z_1 = \frac{V}{n_Q}$$

- extends what we already did, to find free energies like  $F$ ,  $G$  and  $H$  for a Canonical treatment of the ideal gas. *All ensembles* are supposed to give the same results for these in the thermodynamic limit!
- calculates the Sackur-Tetrode entropy  $S(T, V, N)$  and goes over the notion of *entropy of mixing* and the *Gibbs paradox*.
- discusses as an example, the heat capacity of the diatomic ideal gas. B&B does quantitatively what we are asked to do in G&T Problem 6.46 this week ... find  $Z = Z_{trans}Z_{rot}$  for a diatomic molecule that can rotate. (Not assigned, but G&T Problem 6.47 talks about the specific heat of molecules that can vibrate too ... so we'd begin with  $Z = Z_{trans}Z_{rot}Z_{vib}$  where  $Z_{vib}$  is derived as G&T Eq. (4.128) and B&B Eq. (20.28).

- ***Simulation techniques:*** In our G&T reading, we learn
- a tiny bit about *micro canonical simulations*, which will be more relevant when we do magnetic systems (because other things than energy are of interest)
- *demon thermometer simulations*
- *canonical simulations*, which could be done by solving  $F = ma$  (molecular dynamics) but are easier to do via *Monte Carlo (MC)*. A description of the standard Metropolis algorithm is given, which involves the very useful accept/reject paradigm that will establish thermal equilibrium at temperature  $T$ .