

Physics 114 Statistical Mechanics Spring 2018
Seminar 1

Break provided by: Amy

Overview:

Things fall. -Isaac Newton

Things fall apart. -William Butler Yeats; Chinua Achebe

We will revisit what you know of thermodynamics this week ... reminding ourselves of the 0th law (thermometers work) and 1st law (increase in energy of a system = work on it + heat into it), and allude to the 2nd law (things fall apart). Thermo is an extremely important field which brought about the industrialization of society and revolutionized transportation, power generation, chemical engineering ... It is a wholly self-consistent set of ideas and equations based on experimental definitions and multivariable calculus. Weirdly, thermo is largely immune to microscopic details of the gasses, liquids, solids and plasmas that it describes.

Why is thermodynamics valid? The answer lies in *statistical mechanics* which treats large systems comprised of small objects. Stat mech rests on some ridiculously simple postulates about these systems, like ergodicity and equal a priori probabilities. “From micro to macro” is the first theme in one of our texts, Gould and Tobochnik (G&T). The implication is that there are many microscopic details that we *do not need to know*, to calculate the macroscopic properties correctly. (Details that we *do* need include whether the objects obey classical or quantum mechanics, whether they are relativistic or not, and if there are constraints on what they can do ... like from walls, insulation, etc.) Thanks to the laws of probability, *stat mech provides the definitions and supports the equations that comprise thermodynamics*. Moreover, non-equilibrium stat mech takes us beyond thermo, to systems whose macroscopic properties evolve in time. Stat mech even applies to non-physics systems like biological and economic ones where the “laws of motion” are not Newtonian or the Schrodinger Equation, but particular to those disciplines. *Stat mech rocks!*

Our first job this week will be to go over the organization of the seminar and decide how you want it to run, how you want your grade to be determined, and so on. Usually, I like to have both presentations and problem-solving, both analytical and numerical work. As you can tell from the long “Concept checklist” part of this assignment, I feel compelled to *summarize in writing* the major ideas each week. Hopefully, this goes a long way toward helping you organize your learning, study for tests, ... I also like to take a little time during seminar to summarize a major idea - and even more importantly - to have you to comment on and ask questions about what you’ve read.

Required Reading:

G&T sections

- chapter 1
- from chapter 2, sections 2.1 - 2.11

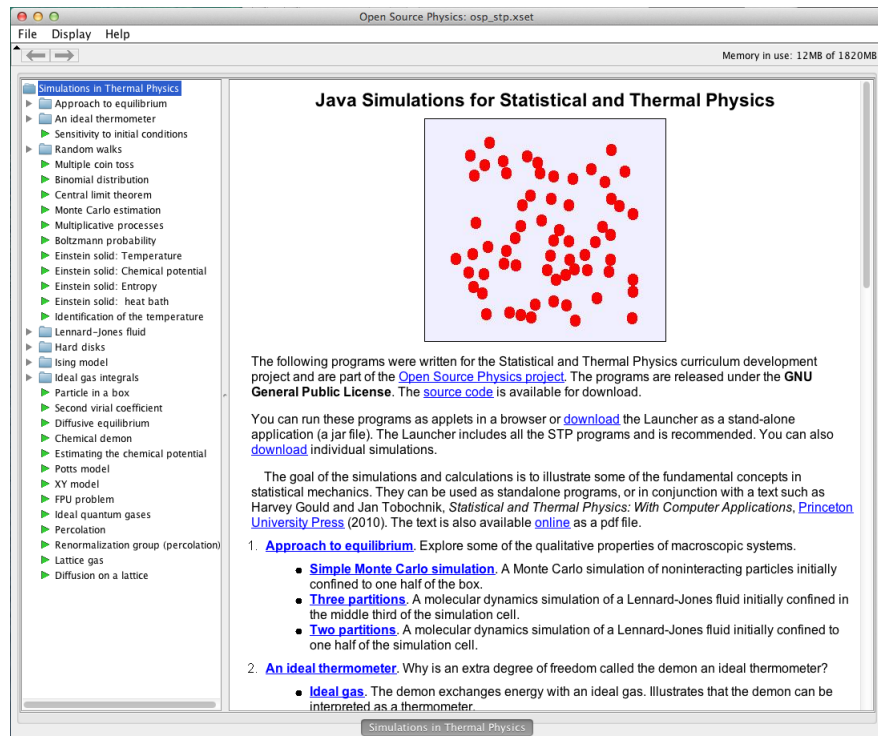
B&B sections

- from chapter 1, sections 1.1-1.3
- Appendices C.6 and C.7
- chapter 2
- chapter 11
- chapter 12

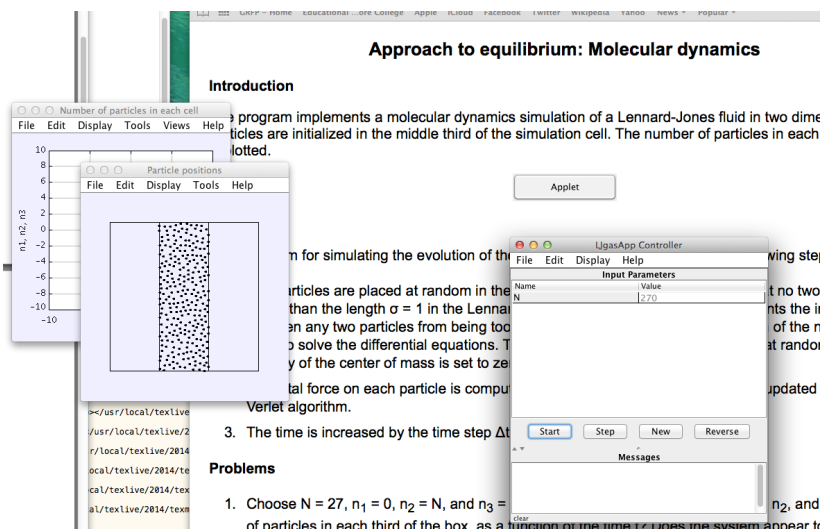
Computational Preparation: Be sure that whatever computer you typically use is able to run the simulations that go with G&T . One thing that might work is the G&T Statistical and Thermal Physics java website: <http://stp.clarku.edu/simulations/> . Redirect yourself, or else go directly to <https://www.compadre.org/STP/?> . Below is the Compadre page for “STP”

The screenshot shows the homepage of the Statistical and Thermal Physics website. The header features the title "Statistical and Thermal Physics" with a search bar. Below the header is a navigation menu with links: "Browse by Category", "Browse by Topic", "Shared Resources", "Discussion Forums", and "About STP". The main content area is divided into several sections. On the left, there is a "Welcome!" message and a "STP Materials" section with links to "The elusive chemical potential", "Teaching Statistical Physics by Thinking about Models and Algorithms", and "STP Ising 1D Program". On the right, there is a "Statistical and Thermal Physics" section with a list of chapters: "Chapter 1: From Microscopic to Macroscopic Behavior", "Chapter 2: Thermodynamic Concepts and Processes", "Chapter 3: Concepts of Probability", "Chapter 4: The Methodology of Statistical Mechanics", "Chapter 5: Magnetic Systems", "Chapter 6: Many-Particle Systems", "Chapter 7: Chemical Potential and Phase Equilibria", "Chapter 8: Classical Gases and Liquids", and "Chapter 9: Landau Theory and the Renormalization Group". At the bottom right, there is a "Featured Simulation" section with a link to "Molecular Dynamics Demonstration Model".

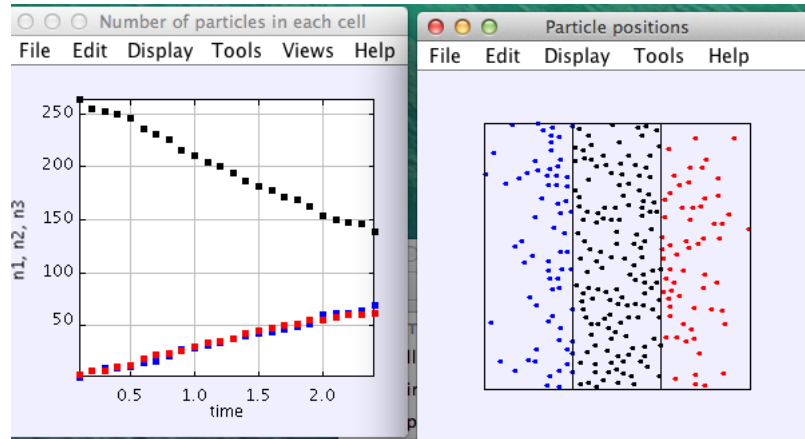
(simulations in thermal physics). If you then follow that page’s link to “Chapter 1: From Microscopic to Macroscopic Behavior” and at the bottom of that next page, select a link saying “Simulations”, you’ll see



The goal is to hit the green arrow next to the “Three partitions” applet and see something like the image below, after hitting “Initialize”, then “Start”.



Hopefully, one way or another you are now running the simulation and taking data! Below is what you should see once you hit “initialize” and start running ApproachToEquilibriumThreePartitions.



Concept checklist from Readings:

- Time has a direction for a large system ... “things fall apart”
- Energy naturally goes from a useful (able to do work) to a less useful (thermal energy) form.
- We might be tempted to say that energy goes from the form of work to the form of heat. But that’s wrong ... why?
- How is heat, Q different from energy, E ? How is it different from temperature, T ?
- *Notation alert:* B&B uses U for energy, whereas G&T uses E . Look for other such notational differences as we read. Sigh.
- *System* and its surroundings: System can be open or closed; walls insulating or conducting.
- *Microstates* and *macrostates*: Their definition depends on what we are modeling (gas? magnet? black body?) For a classical, monatomic gas of N atoms, the macrostate for a closed system is characterized by three variables: P , V and T while a microstate has $6N$ variables, the positions and momenta of the atoms. N is large (on the computer, 10^6 and in real life, 10^{23}).
- Thermodynamic variables can be *intensive* or *extensive*.
- *Entropy* (whose rigorous definition is postponed for a future week) is something that increases in all real-world processes. It is also something that labels a macrostate. It is larger for macrostates that correspond to larger numbers of microstates.
- Simulations: A *molecular dynamics* simulation solves $\vec{F} = d\vec{p}_i/dt$ numerically, to find $\vec{r}_i(t)$ and $\vec{v}_i(t)$ for $i = 1, 2, \dots, N$ particles.

- Systems *spontaneously approach equilibrium*. In equilibrium we observe a mean, or average, value of a macroscopic quantity. Yes, this definition is somewhat circular ... the system is in equilibrium when its macroscopic properties don't change and are described by the laws of (equilibrium) thermodynamics.
- Macroscopic properties are often called *functions of state* (or variables of state or state variables). Functions of state change if we change *parameters* of the system, and their change is always described as an exact differential. Thus, they are independent of the path in parameter space, depending only on the starting and ending values of the parameters.
- Equilibrium is not the same thing as being static. There are *equilibrium fluctuations* in equilibrium mean values. E.g. the simulation `ApproachToEquilibriumParticlesInBox` will show us a mean number on the left side, \bar{n} , but also a variance, $\sigma^2 = \overline{(n - \bar{n})^2}$.
This week, I just want you to get a flavor for means and variances (which measure fluctuations around the mean). In later weeks, we will calculate these theoretically, and estimate them with simulations. We will also see how means and variances scale with system size, temperature, ...
- The postulates of *history independence* and *equal a priori probabilities*.
- Pressure arises from microscopic collisions of particles with walls.

$$P = F/A = \frac{1}{A} \frac{dp_x}{dt}$$

where just for argument's sake, we consider collisions with the area A of a wall in the $y - z$ plane, so the x component of momentum p_x is what changes during the collision.

- *The 0th law of thermodynamics*: One statement is that when two bodies are in thermal equilibrium with a third body, they are in thermal equilibrium with each other. "Thermal equilibrium" means same temperature. Two systems come to the same temperature when they are allowed to exchange energy. Of course, this might not be the original temperature of either system. A thermometer subtracts only a tiny amount of energy from a system. It thus does not change the system's original temperature, but merely reports on it.
- The *ensemble* concept : Many different systems, each one in equilibrium. On the other hand, simulations and experiments often involve just one system advancing in time. The *ergodic hypothesis* lets us equate the two pictures. Thus an ensemble of many equilibrium systems, and one system in equilibrium measured over time will show us the same physics.

- *Sensitivity to initial conditions* suggests that history independence and the ergodic hypothesis apply. As a system evolves, it eventually “forgets” the details of how it began.
- How long does the system take to forget its starting configuration? There will be some *relaxation time* - a time scale for reaching equilibrium. This time depends on the size of the system, the type of contact with its surroundings, etc. You can run G&T programs like `ApproachToEquilibrium...` or `SensitivityToInitialConditions` to get a sense of this for specific cases. G&T section 1.13.1 gives you the case of N particles in a box with two sides, and allows you in Problems 1.2 and Problem 1.10 to calculate the time scale, τ over which the system equilibrates. Here, $\tau = N/2$.
- Models of matter:
 - ideal gas particle
 - particles that interact via a pure repulsion as in G&T Eq. (1.5)
 - the LJ interaction
 - lattice models
- Important constants and their values: k , N_A and R .
- The Kelvin temperature scale. The ideal gas temperature scale is a related idea. We will talk about the “triple point later”.
- A *pressure equation of state* is sometimes written $P = f(T, \rho)$ where $\rho = N/V$; though more casually we just write a function relating P, T, N and V . The ideal gas one is $PV = NkT$ (physicist-preferred) or $PV = nRT$. (chemist-preferred). The van der Waals equation of state corrects some assumptions about idealized gasses.
- An *energy equation of state* looks like $E = f(N, V, T)$. The ideal gas and van der Waals equations are found in G&T section 2.9. Note that energy of an ideal gas does not care about its density, $\rho = N/V$, but for the more realistic gas, energy changes - in the case of Eq. (2.24) it is lower - when density is higher.
- The monatomic ideal gas energy of state $E = \frac{3}{2}NkT$ will be derived later (though you may know it if you’ve already studied kinetic theory of gasses or the equipartition theorem).
- *Processes* take us from one thermodynamic macrostate to another. *Quasi static* means that the system passes through a set of equilibrium states at every stage. *Reversible* means that one can find a practical way to run the process in reverse. For example, any process where friction acts is irreversible. An irreversible process cannot be quasistatic. Subtle? Hell yes.

- The *work done on a gas* undergoing a quasi static process from state 1 to 2 is

$$W_{1 \rightarrow 2} = - \int_{V_1}^{V_2} P dV$$

We can use this equation plus the equation of state for an ideal gas to find the work done on processes that are

- isothermal
- isochoric
- isobaric

- Schroeder (Phys 13) called this kind of work “compression work” which is smart - it reminds you that there are other kinds, though this is the one we most often think of ... gas in cylinder acted upon by piston.
- An *adiabatic* process takes place when the system is insulated from its surroundings, so no heating or cooling happens. For an adiabatic process, $Q = 0$. Careful: this is not saying that the temperature of the system is constant on an adiabat. Rather, work done on the system, W , will change its thermal energy: $W = \Delta E$, and hence its temperature will change. (Exception: The system is undergoing a phase transition.)
- The net work *on* a gas is the negative of area of a cycle like Fig. 2.4 in G&T. If arrows would reversed, sign of work changes. For the work done *by* the gas, again the sign changes.
- *The 1st law of thermo:* $\Delta E = W + Q$.
- For an ideal gas, with the help of the 1st law, we can calculate Q , ΔE , and W for a process as shown on a PV diagram. In particular, know how to do this for the four fundamental kinds of processes: isochores, isobars, isotherms and adiabats.
- For adiabats we derive that $PV^{(f+2)/f} = \text{const}$ where G&T limits us to the case $f = 3$, true for monatomic gasses. Equivalently, $TV^{2/f} = \text{const}$. The abbreviation $\gamma = (f + 2)/f$ is used. The symbol γ is called the “adiabatic index” or “adiabatic constant”.
- We will talk more about *enthalpy*, H , later. For now, note that the definition of this function of state is $H = U + PV$. Schroeder (Phys 13) gave us this memorable way to think about the difference between H and U :



Figure 1.15. To create a rabbit out of nothing and place it on the table, the magician must summon up not only the energy U of the rabbit, but also some additional energy, equal to PV , to push the atmosphere out of the way to make room. The *total* energy required is the *enthalpy*, $H = U + PV$.

- *Heat capacity* is $C = \frac{Q}{\Delta T}$. Different processes might be involved as T changes, yielding different values of C . Two popular cases are isochoric and isobaric changes:

$$C_V \equiv (\partial E / \partial T)_V \quad ; \quad C_P \equiv (\partial H / \partial T)_P$$

The heat capacity for isobaric processes is higher because some of Q goes into work, not increasing E ... hence increasing T . It can be shown that $\gamma = \frac{C_P}{C_V}$. Similarly, $C_P - C_V = R$.

- The *specific heat capacity* is the heat capacity per unit mass of the substance. We often note specific heat as c . Sometimes people want the heat capacity of exactly one mole: this is called the *molar heat capacity*. (It can be a tiny bit confusing to read C_V in B&B and not know if it is for one mole or n moles, but the units give it away.)

Checklist: Math concepts to know this week:

- easy-to-evaluate integrals as in G&T 1.13.2
- given $f(x, y)$, how to take a partial derivative like $(\partial f(x, y) / \partial x)_y$.
- writing a Taylor series for a function
- the definition of an exact differential - for example that the integral of $f(x, y)$ is independent of the path between its endpoints, and that partial differentiation can occur in any order, so $(\partial^2 f / \partial x \partial y) = (\partial^2 f / \partial y \partial x)$

Presentation: No student ones this week

Amy will talk about states, equilibrium and temperature.

Warmup Problems: *Due before seminar, Monday at 5 pm*

Note: In future, you will have just one or two warmup problems!

1. Say which computing languages/environments you feel comfortable using in order to solve numerical problems and make graphs. Which computer languages or environments (if any) could you comfortably use to write a short numerical calculation or simulation?
2. Go to the G&T Statistical and Thermal Physics java website: <http://stp.clarku.edu/simulations/> . Try to run a simulation like **ApproachToEquilibriumThreePartitions**. (This simulation is discussed in Section 1.4 of G&T and I ask you to do a G&T problem using it below.) This might be run as an applet (perhaps necessitating a change in your Java security settings) or as a standalone Java program. Did you succeed? If not, ITS can help ... I hope the answer to this question is ultimately 'yes' :-)
3. **work** G&T Problem 2.7
4. **pressure equation of state** Write the van der Waals equation of state, which is G&T Eq. (2.12) in the form $P = f(T, \rho)$

Regular Problems: *For this first seminar, do as many as you comfortably have time for. For future seminars, we are going to assign problems to pairs of people. Each of you will get a chance to do the “introduction” to a problem, and to do the “solving” in another.*

1: reversible or not? G&T Problem 2.6

2: work G&T Problem 2.8

3: adiabatic and isothermal compression G&T Problem 2.17

4: transfer of heat between two bodies B&B Problem 2.5.

5: isothermal expansion B&B Problem 11.1

6: taking partial derivatives B&B Problem 11.4

7: partial derivatives for a thermal process B&B Problem 12.3

8: work, heat and thermal energy for different thermodynamic processes G&T Problem 2.56

9: free expansion (*part of G&T Problem 2.24, which we will come back to later*) An ideal gas of N particles in a box of volume V_2 is initially confined to only part of the box, with volume $V_1 < V_2$. The remaining volume of the box is empty - a vacuum. The gas has initial temperature T . The gas is then allowed to expand freely to fill the entire volume V_2 . The box is thermally insulated from the outside world. What is the change in energy of the gas? What is the change in temperature of the gas?

10: modeling a bicycle pump G&T Problems 1.24 and 2.9

11: approach to equilibrium with three and two partitions G&T problem 1.1 parts (a) - (f)

If the VideoCapture feature of these applications doesn't work for you (It doesn't for me.) you can make a video some other way. E.g. Quicktime player does it on the Mac ... see under the File menu that you can do a New Screen Recording.

12: more about approach to equilibrium G&T problem 1.10

13: $PV = NkT$ A cloud of neutral hydrogen atoms has a mass which is 100 times the mass of our sun, a temperature of $T=50K$ and a number density of 500 atoms per cm^3 . What is its pressure in Pa? What is its volume? (If you are an Astro person, feel free to convert the volume to ly^3 :-)